

Polarized and unpolarised transverse waves, with applications to optical systems

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Outline

Previous lecture:

- The quarter wave plate
- Set up coordinate system suitable for transverse waves
- Jones calculus; matrix formulation of how wave polarization changes when passing through polarizing component
 - Examples: linear polarizer, quarter wave plate, Faraday rotation

This lecture

- Statistical representation of incoherent/unpolarized waves
 - Polarization tensors and Stokes vectors
 - The Poincare sphere
- Muller calculus; matrix formulation for the transmission of partially polarized waves

Incoherent/unpolarised

- Many sources of electromagnetic radiation are not coherent
 - they do not radiate perfect harmonic oscillations (not sinusoidal wave)
 - over short time scales the oscillations look harmonic
 - but over longer periods the wave look incoherent, or even stochastic
 - such waves are often referred to as unpolarised
- To model such waves we will consider the electric field to be a stochastic process, i.e. it has
 - an average: $\langle E^{\alpha}(t, \mathbf{x}) \rangle$
 - a variance: $\langle E^{\alpha}(t,\mathbf{x}) [E^{\beta}(t,\mathbf{x})]^* \rangle$
 - a covariance: $\langle E^{\alpha}(t,\mathbf{x}) [E^{\beta}(t+s,\mathbf{x}+\mathbf{y})]^* \rangle$
- In this chapter we will focus on the variance, here called the intensity tensor

$$I^{\alpha\beta} = \langle E^{\alpha}(t,\mathbf{x}) [E^{\beta}(t,\mathbf{x})]^* \rangle$$

and the polarization tensor (where $e_M = \mathbf{E} / |\mathbf{E}|$ is the polarization vector)

$$p^{\alpha\beta} = \langle e_M^{\alpha}(t,\mathbf{x})[e_M^{\beta}(t,\mathbf{x})]^* \rangle$$

Representations for the polarization tensor

- The polarisation tensor: $p^{\alpha\beta} = \langle e^{\alpha^*}e^{\beta} \rangle$
 - It has four complex components
 - But $p^{\alpha\beta}$ is constructed from a normalised vector
 - It's components are not all independent!
 - What are the possible forms for $p^{\alpha\beta}$?
- Some restrictions/properties:
 - The tensor has unit trace: $tr(\mathbf{p}) = p^{\alpha\alpha} = \langle e^{\alpha^*}e^{\alpha} \rangle = \langle |\mathbf{e}| \rangle = \langle 1 \rangle = 1$
 - It's hermitian: $(p^{\alpha\beta})^* = \langle e^{\alpha}e^{\beta^*} \rangle = \langle e^{\beta^*}e^{\alpha} \rangle = p^{\beta\alpha}$
- The polarisation tensor is described by three real parameter $\{q,u,v\}$

$$p^{\alpha\beta} = \frac{1}{2} \begin{pmatrix} \begin{bmatrix} 1 & 0 \\ 0 & 1 \end{bmatrix} + q \begin{bmatrix} 1 & 0 \\ 0 & -1 \end{bmatrix} + u \begin{bmatrix} 0 & 1 \\ 1 & 0 \end{bmatrix} + v \begin{bmatrix} 0 & -i \\ i & 0 \end{bmatrix} \end{pmatrix}$$

- Note: the three last ones are the Pauli matrixes.
- These form a basis for an 2x2 hermitian matrix.

Examples

- For example consider:
 - linearly polarised waves in the horisontal plane $e_M^{\alpha} = [1, 0]$

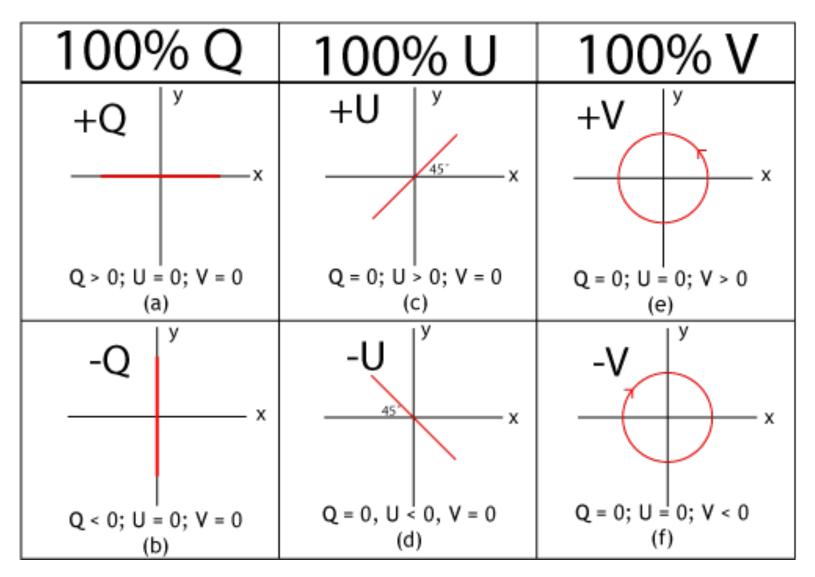
- rotate linear polarization by 45°, $e_M^{\alpha} = [1, 1]2^{1/2}$

$$p^{\alpha\beta} = e_M^{\alpha*} e_M^{\beta} = \frac{1}{2} \begin{bmatrix} 1 \\ 1 \end{bmatrix} \begin{bmatrix} 1 \\ 1 \end{bmatrix} = \frac{1}{2} \begin{bmatrix} 1 \\ 1 \end{bmatrix} = \frac{1}{2} \begin{bmatrix} 1 \\ 0 \end{bmatrix} + 1 \begin{bmatrix} 0 \\ 1 \end{bmatrix} + 1 \begin{bmatrix} 0 \\ 1 \end{bmatrix}$$
$$\rightarrow \{q, u, v\} = \{0, 1, 0\}$$

- right hand circularly polarised waves, $e_M^{\alpha} = [1, -i] / 2^{1/2}$

$$p^{\alpha\beta} = e_M^{\alpha*} e_M^{\beta} = \frac{1}{2} \begin{bmatrix} 1 \\ i \end{bmatrix} \begin{bmatrix} 1 \\ -i \end{bmatrix} = \frac{1}{2} \begin{bmatrix} 1 \\ i \end{bmatrix} = \frac{1}{2} \left(\begin{bmatrix} 1 & 0 \\ 0 & 1 \end{bmatrix} + 1 \begin{bmatrix} 0 & -i \\ i & 0 \end{bmatrix} \right)$$
$$\rightarrow \{q, u, v\} = \{0, 0, 1\}$$

Table of ideal polarisations



By Dan Moulton - http://en.wikipedia.org/wiki/Image:Side2.png, CC BY-SA 3.0, https://commons.wikimedia.org/w/index.php?curid=3319458

The Stokes vector

• Using unit- and Pauli-matrixes, we define $\tau_{j}^{\alpha\beta}$ as:

$$\tau_1^{\alpha\beta} = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}, \ \tau_2^{\alpha\beta} = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}, \ \tau_3^{\alpha\beta} = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \ \tau_4^{\alpha\beta} = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}$$

These can also be used to express the intensity tensor:

$$I^{\alpha\beta} = \frac{1}{2} \begin{bmatrix} I \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} + Q \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} + U \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} + V \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix} \end{bmatrix} = \frac{1}{2} \begin{bmatrix} I + Q & U - iV \\ U + iV & I - Q \end{bmatrix}$$

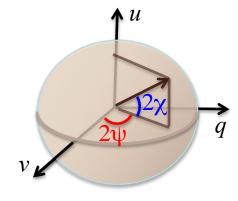
- The four parameters are called the Stokes parameter { I, Q, U, V }
- The Stokes vector is similarly defined as $S_A\big|_{A=\{1,2,3,4\}}= \big[I,Q,U,V\big]$
- Using index notation the intensity matrix and the Stokes vector are related by:

$$I^{\alpha\beta} = \frac{1}{2} \tau_A^{\alpha\beta} S_A$$
 with inverse: $S_A = \tau_A^{\alpha\beta} I^{\alpha\beta}$

- The matrixes $\tau_j^{\alpha\beta}$, defines a transformation between hermitian 2x2 matrixes and real 4-vectors

Poincare sphere

- Define the degree of polarisation: $r = \sqrt{q^2 + u^2 + v^2}$
- Consider the normalised vector $\{q/r, u/r, v/r\}$ (polarised fraction)
 - since this vector is real and normalised it will represent points on a sphere, the so called *Poincare sphere*
- Thus, any transverse wave field can be described by
 - a point on the Poincare sphere
 - a degree of polarization, r



Poincare sphere

- A polarizing element induces a motion on the sphere
 - e.g. passing though a birefringent crystal traces a circle :
 - Birefringence rotates in a vertical plane
 - Faraday rotation rotates in a horisontal plane
 - Here is may be useful to polar coordinates (ψ,χ) .

The polarization tensor for unpolarized waves (1)

- What are the Stokes parameters for unpolarised waves?
 - Let the $e_M^{\ l}$ and $e_M^{\ 2}$ be independent stochastic variable

$$p^{\alpha\beta} = <\begin{pmatrix} e_M^1 \\ e_M^2 \end{pmatrix} \begin{pmatrix} e_M^1 & e_M^2 \end{pmatrix}^* > = \begin{bmatrix} & \\ & \end{bmatrix}$$

- Since $e_M^{\ I}$ and $e_M^{\ 2}$ are uncorrelated the offdiagonal term vanish

$$p^{\alpha\beta} = \begin{bmatrix} < |e_M^1|^2 > & 0 \\ 0 & < |e_M^2|^2 > \end{bmatrix}$$

- The vector \mathbf{e}_M is normalised: $\left|e_M^1\right|^2 + \left|e_M^2\right|^2 = 1$
- By symmetry (no physical difference between $e_M^{\ l}$ and $e_M^{\ 2}$)

$$\left| e_M^1 \right|^2 = \left| e_M^2 \right|^2 = 1/2$$

the polarization tensor then reads

$$p^{\alpha\beta} = \frac{1}{2} \begin{bmatrix} 1 & 0 \\ 0 & 1 \end{bmatrix}$$

- i.e. unpolarised have $\{q,u,v\}=\{0,0,0\}$!

The polarization tensor for unpolarized waves (2)



- Alternative derivation; polarization vector for unpolarized waves
 - Note first that the polarization vector is normalised

$$\left|\mathbf{e}_{M}\right|^{2} = \left|e_{M}^{1}\right|^{2} + \left|e_{M}^{2}\right|^{2} = 1 \sim \cos^{2}(\theta) + \sin^{2}(\theta)$$

- the polarization is complex and stochastic: $\begin{pmatrix} e_M^1 \\ e^2 \end{pmatrix} = \begin{pmatrix} e^{i\phi_1} \cos(\theta) \\ e^{i\phi_2} \sin(\theta) \end{pmatrix}$
 - where θ , ϕ_1 and ϕ_2 are uniformly distributed in $[0,2\pi]$
- The corresponding polarization tensor

$$p^{\alpha\beta} = <\begin{pmatrix} e_M^1 \\ e_M^2 \end{pmatrix} \begin{pmatrix} e_M^1 & e_M^2 \end{pmatrix}^* > = <\begin{pmatrix} e^{i\phi_1 - i\phi_1} \cos(\theta) \cos(\theta) & e^{i\phi_1 - i\phi_2} \cos(\theta) \sin(\theta) \\ e^{i\phi_2 - i\phi_1} \sin(\theta) \cos(\theta) & e^{i\phi_2 - i\phi_2} \sin(\theta) \sin(\theta) \end{pmatrix} >$$

- here the average is over the three random variables θ , ϕ_1 and ϕ_2

$$p^{\alpha\beta} = \frac{1}{(2\pi)^3} \int_{0}^{2\pi} d\theta \int_{0}^{2\pi} d\phi_1 \int_{0}^{2\pi} d\phi_2 \begin{pmatrix} \cos^2(\theta) & e^{i\phi_1 - i\phi_2} \cos(\theta) \sin(\theta) \\ e^{i\phi_2 - i\phi_1} \sin(\theta) \cos(\theta) & \sin^2(\theta) \end{pmatrix} = \frac{1}{2} \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}$$

- i.e. unpolarised have $\{q,u,v\}=\{0,0,0\}$!

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- Muller calculus; matrix formulation for the transmission of partially polarized waves
 - General theory for dispersive media
 - Optical components

Weakly anisotropic media

- In weakly anisotropic media the wave equation can be rewritten on a form suitable for studying the wave polarisation.
- Write the weakly anisotropic transverse response as

$$K^{\alpha\beta} = n_0^2 \delta^{\alpha\beta} + \Delta K^{\alpha\beta}$$

- where $\Delta K^{\alpha\beta}$ is a small perturbation
- The wave equation

$$(n^2 \delta^{\alpha\beta} - K^{\alpha\beta}) E^{\beta} = 0 \implies (n^2 - n_0^2) E^{\alpha} = \Delta K^{\alpha\beta} E^{\alpha}$$

- when ΔK_{ij} is a small, the 1st order dispersion relation reads: $n^2 \approx n_0^2$
- the left hand side can then be expanded to give

$$n^{2} - n_{0}^{2} = (n - n_{0})(n + n_{0}) = (n - n_{0})n_{0} \left[2 + \frac{(n - n_{0})}{n_{0}} \right] \approx 2n_{0}(n - n_{0})$$

$$2n_{0}(n - n_{0})E^{\alpha} \approx \Delta K^{\alpha\beta}E^{\alpha}$$

The wave equation in Jones calculus

• Inverse Fourier transform, when $k_0 = \omega n_0/c$:

$$2k_0(k-k_0)E^{\alpha} \approx \frac{\omega^2}{c^2}\Delta K^{\alpha\beta}E^{\alpha} \Leftrightarrow 2k_0(-i\frac{\partial}{\partial x}-k_0)E^{\alpha} \approx \frac{\omega^2}{c^2}\Delta K^{\alpha\beta}E^{\alpha}$$

• Factor our the eikonal with wave number k_0 :

$$E^{\alpha} = E_0^{\alpha}(x) \exp(ik_0 x)$$

The wave equation can then be simplified

$$(-i\frac{\partial}{\partial x} - k_0)E_0^{\alpha}(x)\exp(ik_0x) \approx \frac{\omega^2}{2k_0c^2}\Delta K^{\alpha\beta}E_0^{\alpha}(x)\exp(ik_0x)$$

$$\frac{dE_0^{\alpha}}{dx} \approx i \frac{\omega}{2n_0 c} \Delta K^{\alpha\beta} E_0^{\beta} \qquad E_{out}^{\alpha} = J^{\alpha\beta} E_{in}^{\beta} \; \; ; \; J^{\alpha\beta} \approx \exp \left[i \frac{\omega}{2n_0 c} \Delta K^{\alpha\beta} x \right]$$

The differential transfer equation in the Jones calculus! (We will use this relation in the next lecture)

The wave equation as an ODE

Wave equation for the intensity tensor:

$$\frac{dI^{\alpha\beta}}{dx} = \frac{d}{dx} < E^{\alpha}E^{\beta^*} > = \dots = \frac{i\omega}{2cn_0} \left(\Delta K^{\alpha\rho}\delta^{\beta\sigma} - \Delta K^{\beta\sigma^*}\delta^{\alpha\rho}\right)I^{\rho\sigma}$$
from prev. page:
$$\frac{dE^{\alpha}_0}{dx} \approx i\frac{\omega}{2n_0c}\Delta K^{\alpha\beta}E^{\beta}_0$$

• Rewrite it in terms of the Stokes vector: $S_A = \tau_A^{\alpha\beta} I^{\alpha\beta}$

$$\frac{dS_{A}}{dx} = (\rho_{AB} - \mu_{AB})S_{B}$$

$$\begin{cases} \rho_{AB} = \frac{i\omega}{4cn_{0}} \left(\tau_{A}^{\beta\alpha} \Delta K^{H,\alpha\rho} \tau_{B}^{\rho\beta} - \tau_{B}^{\rho\sigma} \Delta K^{H,\sigma\beta} \tau_{A}^{\beta\rho} \right) \\ \mu_{AB} = -\frac{i\omega}{4cn_{0}} \left(\tau_{A}^{\beta\alpha} \Delta K^{A,\alpha\rho} \tau_{B}^{\rho\beta} + \tau_{B}^{\rho\sigma} \Delta K^{A,\sigma\beta} \tau_{A}^{\beta\rho} \right) \end{cases}$$

- we may call this the differential formulation of Muller calculus
- symmetric matrix ho_{AB} describes non-dissipative changes in polarization
- and the antisymmetric matrix μ_{AB} describes dissipation (absorption)

The wave equation as an ODE

• The ODE for S_A has the analytic solution (cmp to the ODE y'=ky)

$$S_A(x) = \left[\delta_{AB} + \left(\rho_{AB} - \mu_{AB}\right)x + 1/2\left(\rho_{AC} - \mu_{AC}\right)\left(\rho_{CB} - \mu_{CB}\right)x^2 + ...\right]S_B(0)$$

cmp with Taylor series for exponential

$$S_A(x) = M_{AB}S_B(0)$$
, where $M_{AB} = \exp[(\rho_{AB} - \mu_{AB})x]$

- Here M_{AR} is called the *Muller matrix*
 - $-M_{AB}$ represents entire optical components / systems
 - This is a component based Muller calculus

Examples of Muller matrixes

Some common Muller matrixes:

Linear polarizer (Horizontal Transmission)

Linear polarizer (45° transmission)

Quarter wave plate (fast axis horizontal)

$$M_{AB}^{Q,H} = \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 0 & -1 \\ 0 & 0 & 1 & 0 \end{bmatrix}$$

Attenuating filter (30% Transmission)

(fast axis horizontal) (30% Transmission)
$$M_{AB}^{Q,H} = \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 0 & -1 \\ 0 & 0 & 1 & 0 \end{bmatrix} \qquad M_{AB}^{Att}(0.3) = 0.3 \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

What is the Muller matrix for Faraday rotation?

Examples of Muller matrixes

- In optics it is common to connect a series of optical elements
- consider a system with:
 - a linear polarizer and
 - a quarter wave plate

$$S_A^{out} = M_{AB}^{Q,H} M_{BC}^{L,45} S_C^{in}$$

- Insert unpolarised light, $S_A^{in}=[1,0,0,0]$
 - Step 1: Linear polariser transmit linearly polarised light

$$S^{step1} = M_{BC}^{L,45} \begin{bmatrix} 1 & 0 & 0 & 0 \end{bmatrix}^T = \begin{bmatrix} 1 & 0 & -1 & 0 \end{bmatrix}^T$$

Step 2: Quarter wave plate transmit circularly polarised light

$$S^{out} = M^{Q,H} \begin{bmatrix} 1 & 0 & -1 & 0 \end{bmatrix}^T = \begin{bmatrix} 1 & 0 & 0 & -1 \end{bmatrix}^T$$