

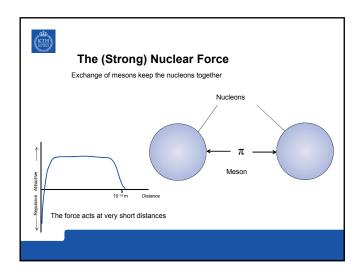


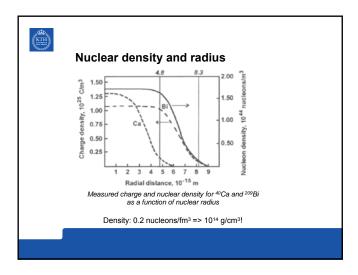
The Nucleus

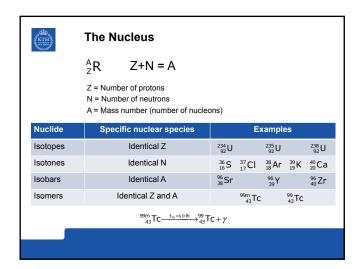


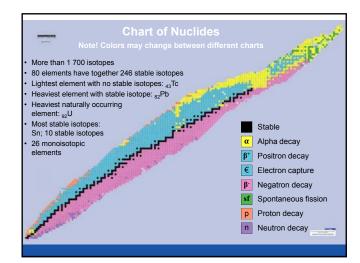
- Building blocks: Protons and neutrons
- Forces: Electromagnetic forces and the Strong Nuclear Force

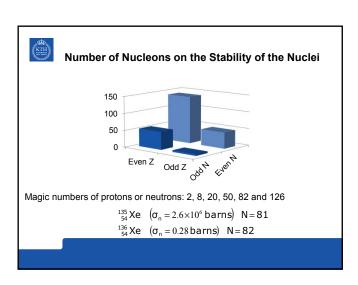
Property	Proton	Neutron
Mass	1.673×10 ⁻²⁴ g	1.675×10 ⁻²⁴ g
Charge	+1	0
Spin	s = ½	S = ½

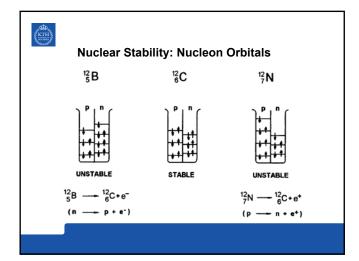














Mass Defect (ΔM) and Mass Excess (δ_A)

 $\Delta M_A = M_A - Z M_p - N M_n$ $M_A = Mass of atom$ $M_H = Mass of Proton (hydrogen)$ $M_n = Mass of Neutron$

Deuterium, $\begin{aligned} &M_p + M_n = 1.007\ 825 + 1.008\ 665 \\ ^2_1H = 2.016\ 490\ u \\ &M_A = 2.014\ 102\ u \\ => &\Delta M_A = -0.002\ 388\ u \end{aligned}$

All stable isotopes have negative mass defect, ΔM_A

Mass excess: $\delta_A = M_A - A$ (sometimes used in tables, no practical use)

M = atomic mass unit, measured in u.



Binding energy

 $\Delta E = \Delta mc^2$

"Nuclear Heat of formation"

 $\Delta m = M_A - (Z m_p + N m_n)$ c = 3 × 10⁸ m/s

Nucleus: 5-10 MeV / nucleon (5-10 \times 10¹¹ J/mol)

Covalent bond: 4.4 × 10⁵ J/mol



Isotope effects

Due to the difference in nucleons there are very small differences between two isotope's

- Freezing point
- Boiling point
- Density
- · Heat of vaporization
- Viscosity
- Surface tension
- · Optical emission spectra



Isotope effects

Replacing ¹H with D (deuterium, ²H) increases the mass 100%

Replacing $^{12}\mbox{C}$ with $^{13}\mbox{C}$ increases the mass 8%

- A reaction involving C-H bond is typically 6-10 times faster than that for a C-D bond
- A reaction involving ¹²C-H bond is 1.04 times faster than that for a ¹³C-H bond



Isotope separation

- i. Equilibrium processes (light elements)
- ii. Rate processes

Multi-stage processes (for instance distillation)

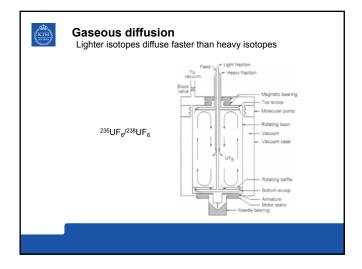
Chemical exchange

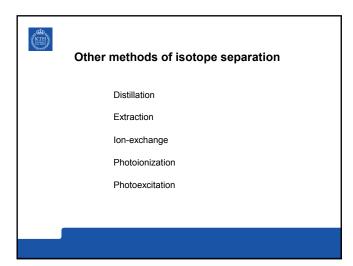
Electrolysis

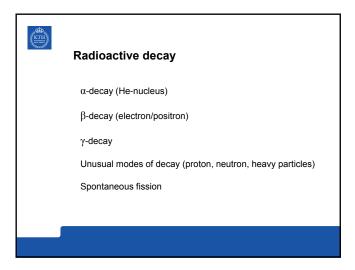
Gaseous diffusion

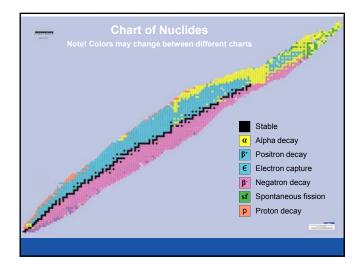
Electromagnetic separation

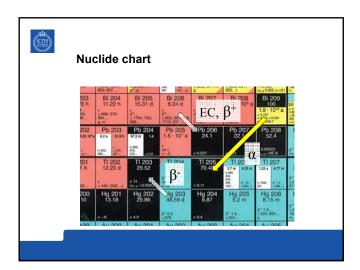
Gas centrifugation

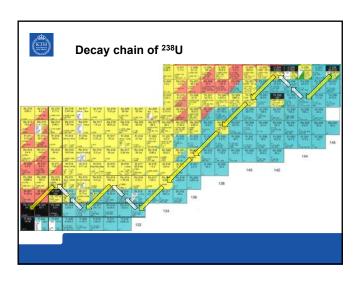


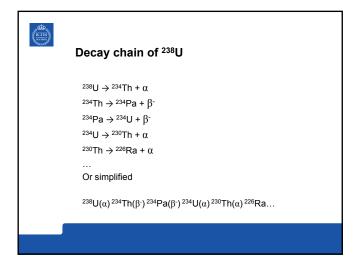


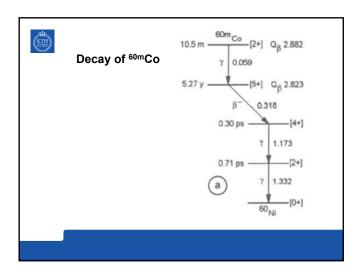


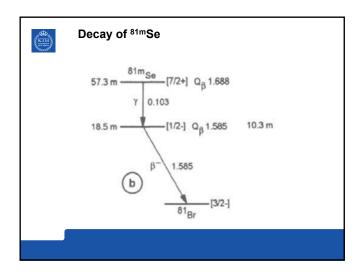


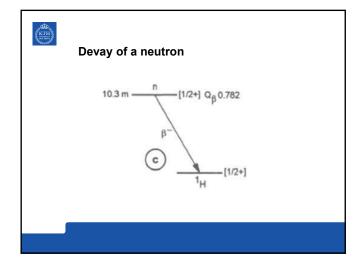


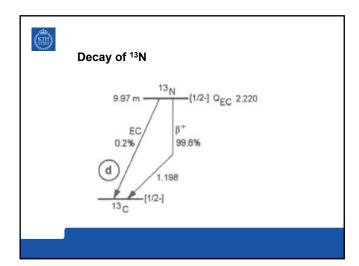


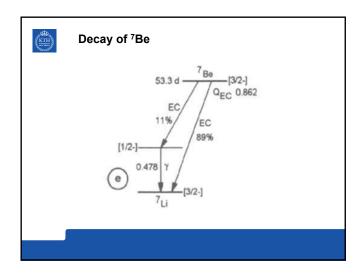


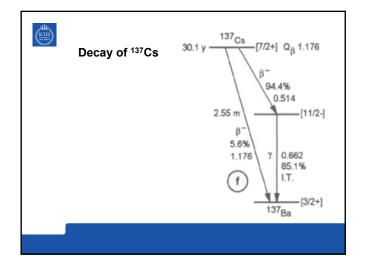


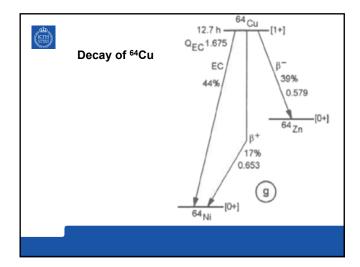


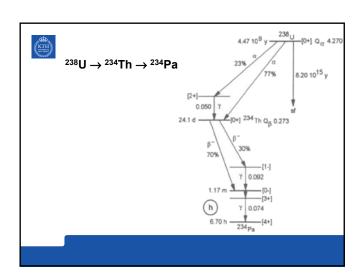


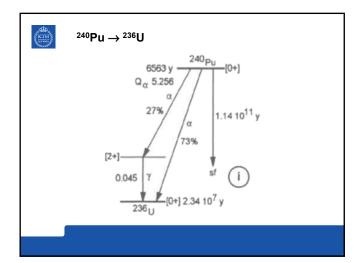












(KTH)

Conservation laws

Nuclear Reaction: $X_1 + X_2 \rightarrow X_3 + X_4$

Energy (mass): $E_1 + E_2 = E_3 + E_4$

Linear momentum: p=mv $p_1 + p_2 = p_3 + p_4$

Charge: $Z_1 + Z_2 = Z_3 + Z_4$

Mass number: $A_1 + A_2 = A_3 + A_4$



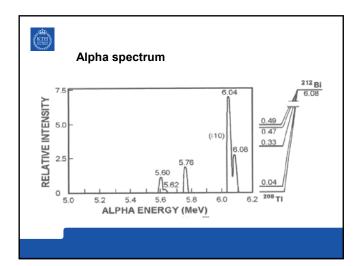
α-decay

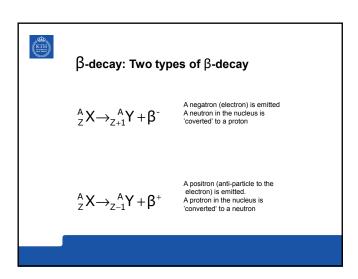
$$_{z}^{A}X \rightarrow _{z-2}^{A-4}Y + _{2}^{4}He$$

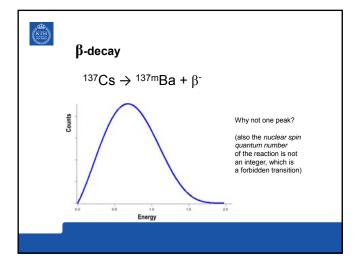
$$^{238}\text{U} \rightarrow ^{234}\text{Th} + ^{4}\text{He}$$

$$^{238}U \rightarrow ^{234}Th + \alpha$$

(KTH)	Decay energy (Q-value)
	E=mc ²
	1 u = 1/6.022×10 ²³ = 1.66×10 ⁻²⁴ g
	$c^2 = 8.99 \times 10^{16} \text{ m}^2/\text{s}^2$
	1 J = 6.24×10 ¹² MeV
	E = 1.66×10 ⁻²⁴ * 8.99×10 ¹⁶ * 6.24×10 ¹² = 931.5 MeV/u
	Q(MeV) = -931.5 ΔM (u)
	$Q_{rr} = -931.5 (M_{Z-2} + M_{He} - M_Z)$
	$Q_{\alpha} > 0$ if $(M_{Z-2} + M_{He} - M_Z) < 0$
	Q _α > 0 => Spontaneous decay
	For α-particles the Q-value is 2–10 MeV









β-decay continued

 \Rightarrow Another particle is emitted: a neutrino (v)

The neutrino has no charge and very small or no mass and does not interact readily with matter

$$^{137}\text{Cs} \rightarrow ^{137}\text{mBa} + \beta^{-} + \overline{\nu}$$

 ν is an anti-neutrino, emitted in a $\beta^\text{-}\text{-decay}$ $\overline{\nu}$ is a neutrino, emitted in a $\beta^\text{+}\text{-decay}$



Energy of β^- decay



$$_{z}^{A}X\rightarrow_{z+1}^{A}Y+\beta^{-}+\stackrel{-}{\nu}$$

The formed Y has Z orbit electrons and must capture one electron from the surroundings.

=> The mass of the $\beta^{\text{-}}\text{-particle}$ shall thus $\underline{\textbf{not}}$ to be included when calculating the energy of the decay.

$$Q_{\beta^-} = -931.513 (M_{Z+1} - M_Z)$$

Example: $n \to H + \beta$

 Q_{β^-} = -931.513(1.007825 - 1.008665) = 0,782 MeV



Energy of β^+ decay



$$_{z}^{A}X \rightarrow _{z-1}^{A}Y^{-} + \beta^{+} + \nu \rightarrow _{z-1}^{A}Y + e^{-} + \beta^{+} + \nu$$

The formed Y has now one extra orbit electron which it must loose.

=> Both emitting a $\beta^{+}\text{-particle}$ and loosing an electron must be included when calculating the energy of the decay.

$$Q_{\beta^-} = -931.513 (M_{Z-1} + 2 M_e - M_Z)$$

Example: $^{13}_{7}N \rightarrow ~^{13}_{6}C + \beta^{\scriptscriptstyle +}$

 Q_{β^-} = -931.513(13.003355– 13.005739) + 2* 0.511) = 1,2 MeV



Electron capture

$$_{z}^{A}X \xrightarrow{EC} _{z-1}^{A}Y + \nu$$

An inner shell electron is captured by the nucleus. Energy similar to $\beta^{\text{-}}$ decay.

 $Q_{\beta^-} = -931.513 (M_{Z-1} - M_Z)$



$\gamma\text{-emisson}$

Most α and $\beta\text{-decays}$ do not go all the way to the daughter's ground state.

The remaining energy is released as $\gamma\text{-rays}.$

Isomeric transition

When the meta-stable state is more long-lived

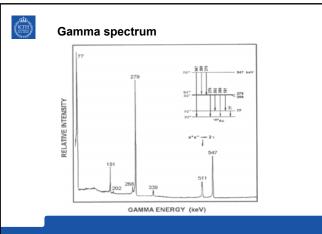
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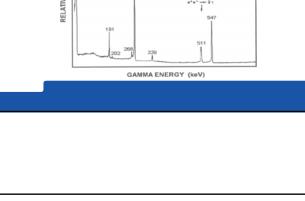
Spontaneous fission

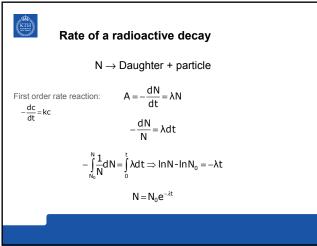
Some heavy radionuclides are so unstable that they undergo spontaneous fission

Rare modes of decay

Proton emission Neutron emission Emission of heavy particles







KTH	Half-life
	$-\int\limits_{N_0}^N \frac{1}{N} dN = \int\limits_0^{t_{y_2}} \lambda dt \Longrightarrow InN - InN_0 = -\lambda t_{y_2}$
	$N = \frac{N_0}{2}$
	$t_{\nu_2} = \frac{ln N_0 - ln \left(\frac{N_0}{2}\right)}{\lambda} = \frac{ln 2}{\lambda}$

KIH	Units
	<u>SI unit:</u> 1 Becquerel (Bq) = 1 decay / s
	<u>Older unit:</u> 1 Curie (Ci) = 3.7 × 10 ¹⁰ Bq
	(1 Ci is approximately the activity of 1 gram ²²⁶ Ra)